

**CALENDAR OF EVENTS** <http://physics.illinois.edu/bluesheet.asp>

**Monday, October 16, ICMT Seminar:** "SU(2)-Invariant Topological Quantum Spin Liquids" Alexander Seidel; 12:00 pm in 190 ESB

**Tuesday, October 17, Astronomy Colloquium: CANCELLED;** 3:45 pm in 134 Astronomy

**Wednesday, October 18, Astrophysics, Gravitation and Cosmology Seminar:** "Measuring the Universe at Microwave, Infrared and Gamma-Ray Wavelengths" Chang Feng; 12:00 pm in 222 Loomis

**Wednesday, October 18, QI/AMO Seminar;** 1:00 pm in 280 MRL

**Wednesday, October 18, Physics Colloquium:** "Super Eddington Black Hole Accretion Flows" Jim Stone; 4:00 pm in 141 Loomis

**Visitors:**

# ICMT Seminar

**Title:** "SU(2)-Invariant Topological Quantum Spin Liquids"

**Speaker:** Alexander Seidel, Washington University in St. Louis

**Date:** Monday, October 16      **Time/Location:** 12:00 pm / 190 ESB

**Abstract:** See here

<https://physics.illinois.edu/calendar/event/10/16/2017/33276964>

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**Visitors:**

# Astrophysics, Gravitation and Cosmology Seminar

**Title:** “Measuring the Universe at Microwave, Infrared and Gamma-Ray Wavelengths”

**Speaker:** Chang Feng (UIUC)

**Date:** Wednesday, October 18

**Time/Location:** 12:00 pm / 222 Loomis

**Abstract:** See here <https://physics.illinois.edu/calendar/event/10/18/2017/33274213>

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**Visitors:**

**Title:** “Super Eddington Black Hole Accretion Flows”

**Speaker:** Jim Stone (Princeton University)

**Date:** Wednesday, October 18

**Time/Location:** 4:00 pm / 141 Loomis

**Abstract:** Super Eddington accretion occurs in many systems, such as the inner regions of quasars and luminous AGN, ultra-luminous X-ray sources (ULXs), and tidal disruption events. Understanding such flows is important not only for interpreting the spectra and variability of these sources, but also to predict the rate of growth of black holes in the early universe, and to quantify energy and momentum feedback into the medium surrounding the black hole, a process likely to be important in controlling galaxy formation in the case of AGN. New results from a study of the magnetohydrodynamics of luminous accretion flows, in which radiation pressure dominates, will be presented. We have developed new numerical methods based on a formal solution of the time-dependent radiation transfer equations to study this regime. Our numerical simulations reveal new effects that require extension of standard thin-disk models. We discuss these results, and their implications for the astrophysics of accreting black holes.